



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.



# TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope Array Collaboration.

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# TELESCOPE ARRAY COLLABORATION

T. Abu-Zayyad<sup>a</sup>, R. Aida<sup>b</sup>, M. Allen<sup>a</sup>, R. Anderson<sup>a</sup>, R. Azuma<sup>c</sup>, E. Barcikowski<sup>a</sup>, J. W. Belz<sup>a</sup>, D. R. Bergman<sup>a</sup>, S. A. Blake<sup>a</sup>, R. Cady<sup>a</sup>, B. G. Cheon<sup>d</sup>, J. Chiba<sup>e</sup>, M. Chikawa<sup>f</sup>, E. J. Cho<sup>d</sup>, W. R. Cho<sup>g</sup>, H. Fujii<sup>h</sup>, T. Fujii<sup>i,\*</sup>, T. Fukuda<sup>c</sup>, M. Fukushima<sup>j</sup>, W. Hanlon<sup>a</sup>, K. Hayashi<sup>c</sup>, Y. Hayashi<sup>c</sup>, N. Hayashida<sup>k</sup>, K. Hibino<sup>l</sup>, K. Hiyama<sup>k</sup>, K. Honda<sup>b</sup>, T. Iguchi<sup>c</sup>, D. Ikeda<sup>a</sup>, K. Ikuta<sup>b</sup>, N. Inoue<sup>m</sup>, T. Ishii<sup>b</sup>, R. Ishimori<sup>i</sup>, H. Ito<sup>d</sup>, D. Ivanov<sup>a</sup>, S. Iwamoto<sup>b</sup>, C. C. H. Juia<sup>a</sup>, K. Kadota<sup>o</sup>, F. Kakimoto<sup>c</sup>, O. Kalashev<sup>k</sup>, T. Kanbe<sup>b</sup>, K. Kasahara<sup>p</sup>, H. Kawai<sup>q</sup>, S. Kawakami<sup>i</sup>, S. Kawanam<sup>m</sup>, E. Kidou<sup>l</sup>, H. B. Kim<sup>d</sup>, H. K. Kim<sup>g</sup>, J. H. Kim<sup>d</sup>, J. H. Kim<sup>a</sup>, K. Kitamoto<sup>f</sup>, S. Kitamura<sup>a</sup>, Y. Kitamura<sup>c</sup>, K. Kobayashi<sup>c</sup>, Y. Kobayashi<sup>c</sup>, Y. Kondo<sup>l</sup>, K. Kuramoto<sup>o</sup>, V. Kuzmin<sup>k</sup>, Y. J. Kwon<sup>g</sup>, J. Lan<sup>a</sup>, S. I. Lim<sup>s</sup>, J. P. Lundquist<sup>a</sup>, S. Machida<sup>c</sup>, K. Martens<sup>l</sup>, T. Matsuda<sup>b</sup>, T. Matsuura<sup>c</sup>, T. Matsuyama<sup>i</sup>, J. N. Matthews<sup>a</sup>, I. Myers<sup>a</sup>, M. Minamino<sup>l</sup>, K. Miyata<sup>o</sup>, Y. Murano<sup>q</sup>, S. Nagatoku<sup>i</sup>, T. Nakamura<sup>v</sup>, S. W. Nam<sup>t</sup>, T. Nonaka<sup>u</sup>, S. Ogio<sup>l</sup>, J. Ogura<sup>c</sup>, M. Ohnishi<sup>l</sup>, H. Ohoka<sup>j</sup>, K. Oki<sup>d</sup>, D. Oku<sup>b</sup>, T. Okuda<sup>l</sup>, M. Ono<sup>u</sup>, A. Oshima<sup>i</sup>, S. Ozawa<sup>p</sup>, I. H. Park<sup>s</sup>, M. S. Pshirkov<sup>w</sup>, D. C. Rodriguez<sup>z</sup>, S. Y. Roh<sup>t</sup>, G. Rubtsov<sup>y</sup>, D. Ryu<sup>u</sup>, H. Sagawa<sup>l</sup>, N. Sakurai<sup>i</sup>, A. L. Sampson<sup>a</sup>, L. M. Scott<sup>a</sup>, P. D. Shah<sup>a</sup>, F. Shibata<sup>b</sup>, T. Shibata<sup>l</sup>, H. Shimodaira<sup>j</sup>, B. K. Shin<sup>d</sup>, J. I. Shin<sup>g</sup>, T. Shirahama<sup>m</sup>, J. D. Smith<sup>a</sup>, P. Sokolsky<sup>a</sup>, T. J. Sonley<sup>a</sup>, R. W. Springer<sup>a</sup>, B. T. Stokes<sup>a</sup>, S. R. Stratton<sup>a,n</sup>, T. A. Stroman<sup>a,l,z</sup>, S. Suzuki<sup>h</sup>, Y. Takahashi<sup>l</sup>, M. Takeda<sup>l</sup>, A. Taketa<sup>x</sup>, M. Takita<sup>l</sup>, Y. Tameda<sup>l</sup>, H. Tanaka<sup>i</sup>, K. Tanaka<sup>y</sup>, M. Tanaka<sup>b</sup>, S. B. Thomas<sup>a</sup>, G. B. Thomas<sup>o</sup>, P. Tinyakov<sup>k,v</sup>, I. Tkachev<sup>k</sup>, H. Tokuno<sup>c</sup>, T. Tomida<sup>b</sup>, S. Troitsky<sup>k</sup>, Y. Tsunesada<sup>c</sup>, K. Tsutsumi<sup>c</sup>, Y. Tsuyuguchi<sup>z</sup>, Y. Uchihori<sup>z</sup>, S. Udo<sup>h</sup>, H. Ukai<sup>b</sup>, G. Vasiloff<sup>z</sup>, Y. Wada<sup>m</sup>, T. Wong<sup>a</sup>, Y. Yamakawa<sup>l</sup>, R. Yamane<sup>i</sup>, H. Yamaoka<sup>h</sup>, K. Yamazaki<sup>i</sup>, J. Yang<sup>s</sup>, Y. Yoneda<sup>i</sup>, S. Yoshida<sup>q</sup>, H. Yoshii<sup>maa</sup>, R. Zollinger<sup>a</sup>, Z. Zundel<sup>a</sup>

<sup>a</sup>University of Utah, High Energy Astrophysics Institute, Salt Lake City, Utah, USA

<sup>b</sup>University of Yamanashi, Interdisciplinary Graduate School of Medicine and Engineering, Kofu, Yamanashi, Japan

<sup>c</sup>Tokyo Institute of Technology, Meguro, Tokyo, Japan

<sup>d</sup>Hanyang University, Seongdong-gu, Seoul, Korea

<sup>e</sup>Tokyo University of Science, Noda, Chiba, Japan

<sup>f</sup>Kinki University, Higashi Osaka, Osaka, Japan

<sup>g</sup>Yonsei University, Seodaemun-gu, Seoul, Korea

<sup>h</sup>Institute of Particle and Nuclear Studies, KEK, Tsukuba, Ibaraki, Japan

<sup>i</sup>Osaka City University, Osaka, Osaka, Japan

<sup>j</sup>Institute for Cosmic Ray Research, University of Tokyo, Kashiwa, Chiba, Japan

<sup>k</sup>Institute for Nuclear Research of the Russian Academy of Sciences, Moscow, Russia

<sup>l</sup>kanagawa University, Yokohama, Kanagawa, Japan

<sup>m</sup>Saitama University, Saitama, Saitama, Japan

<sup>n</sup>Rutgers University, Piscataway, USA

<sup>o</sup>Tokyo City University, Setagaya-ku, Tokyo, Japan

<sup>p</sup>Waseda University, Advanced Research Institute for Science and Engineering, Shinjuku-ku, Tokyo, Japan

<sup>q</sup>Chiba University, Chiba, Chiba, Japan

<sup>r</sup>Chungnam National University, Yuseong-gu, Daejeon, Korea

<sup>s</sup>Ewha Womans University, Seodaemun-gu, Seoul, Korea

<sup>t</sup>Kavli Institute for the Physics and Mathematics of the Universe (WPI), Todai Institutes for Advanced Study, the University of Tokyo, Kashiwa, Chiba, Japan

<sup>u</sup>RIKEN, Wako, Saitama, Japan

<sup>v</sup>Kochi University, Kochi, Kochi, Japan

<sup>w</sup>University Libre de Bruxelles, Brussels, Belgium

<sup>x</sup>Earthquake Research Institute, University of Tokyo, Bunkyo-ku, Tokyo, Japan

<sup>y</sup>Hiroshima City University, Hiroshima, Hiroshima, Japan

<sup>z</sup>National Institute of Radiological Science, Chiba, Chiba, Japan

<sup>aa</sup>Ehime University, Matsuyama, Ehime, Japan



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

## TA detector and data

## Global distributions

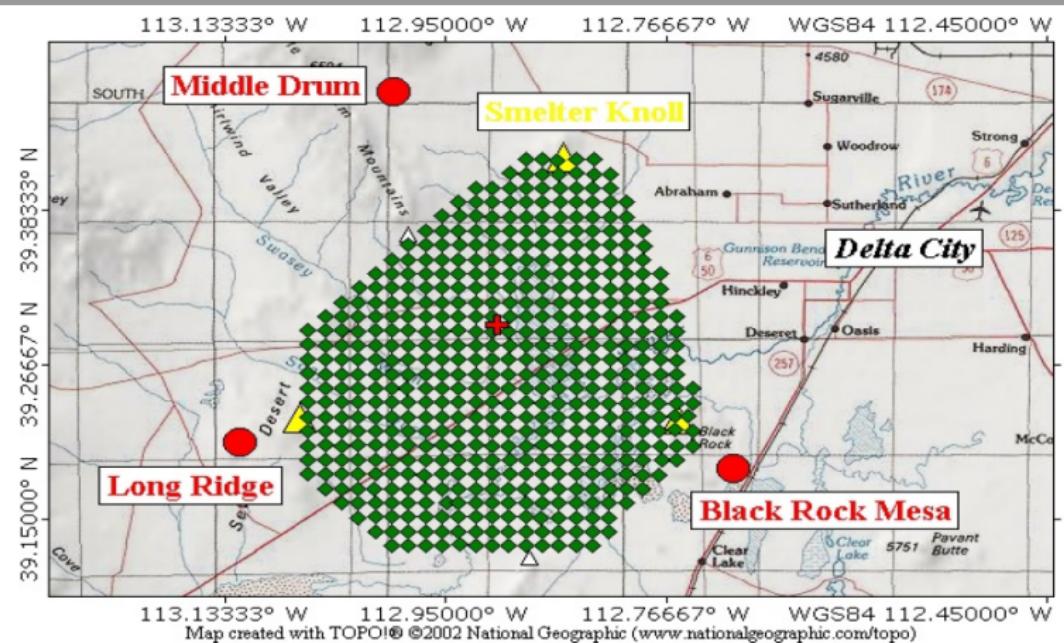
## Point sources

## Hot spot

## Other searches

## Conclusions

# TA HYBRID DETECTOR



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

## TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

- ▶ 507 scintillator detectors covering  $680 \text{ km}^2$
- ▶ 3 fluorescence sites, 38 telescopes
- ▶ SD fully operational from March 2008
- ▶ SD relative size: TA  $\sim 9 \times$  AGASA  $\sim$  PAO/4

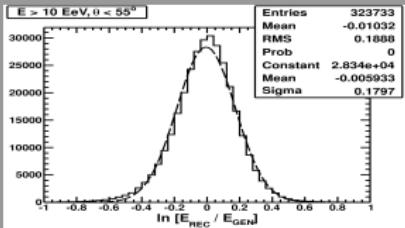
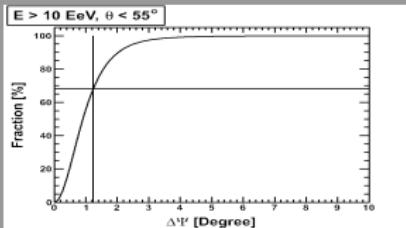
# Anisotropy data set (SD)



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

- ▶ covers the period 12.05.2008 — 11.05.2014 (full 6 years)
- ▶ zenith angle up to  $55^\circ$ , loose border cut
- ▶ geometrical acceptance; exposure  $\sim 7400 \text{ km}^2 \text{ yr sr}$
- ▶ **2560** above 10 EeV
- ▶ **164** above 40 EeV
- ▶ **66** above 57 EeV
- ▶ angular resolution: better than  $1.5^\circ$
- ▶ energy resolution:  $\sim 20\%$



## TA detector and data

## Global distributions

## Point sources

## Hot spot

## Other searches

## Conclusions



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

# GLOBAL DISTRIBUTIONS

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# Comparison with isotropic distribution by KS test

KS p-values:

- ▶ 2560 events with  $E > 10$  EeV

Coords	right ascension	declination
Equatorial:	0.38	0.92
Supergalactic:	0.60	0.35

- ▶ 164 events with  $E > 40$  EeV

Coords	right ascension	declination
Equatorial:	0.075	0.39
Supergalactic:	0.10	0.21

- ▶ 66 events with  $E > 57$  EeV

Coords	right ascension	declination
Equatorial:	0.022	0.16
Supergalactic:	0.10	0.005



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# AUTOCORRELATION FUNCTION



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

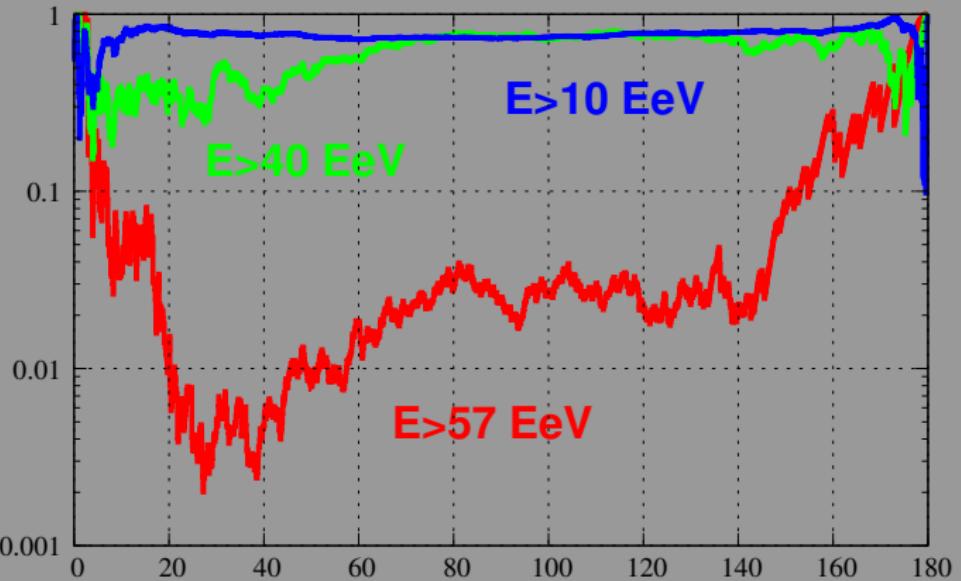
Global distributions

Point sources

Hot spot

Other searches

Conclusions



➡ compatible with isotropy at  $E > 10$  EeV and  
 $E > 40$  EeV, tension at  $E > 57$  EeV



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

# POINT SOURCES

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

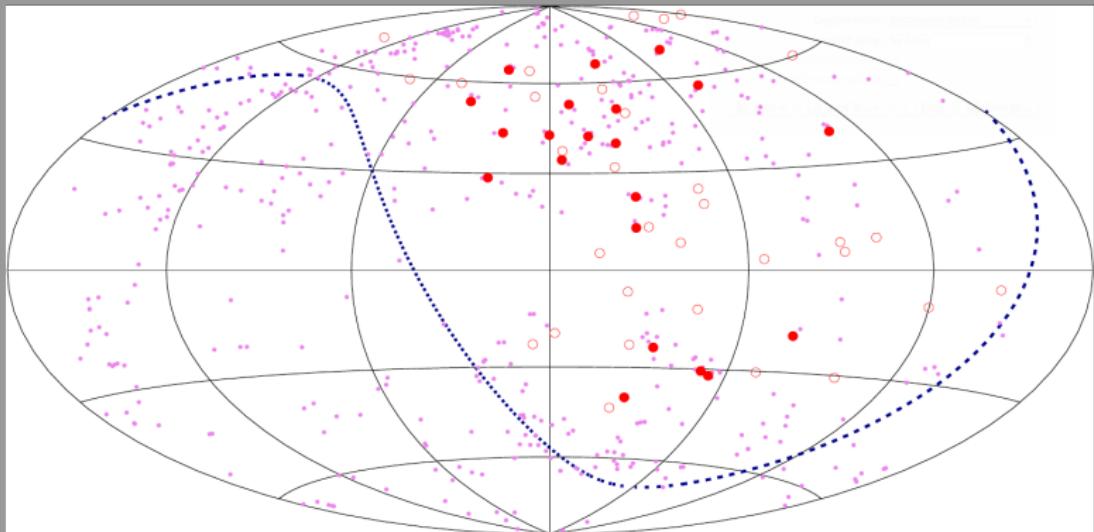
# CORRELATION WITH NEARBY AGN

- ▶ DATA: strict cuts,  $\theta_z < 45^\circ$ ,  $E > 57 \text{ EeV}$ , 6 years: 53 events



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.



dots: AGN from VCV catalog

filled circles: TA events correlating within  $3.1^\circ$

empty circles: non-correlating events

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

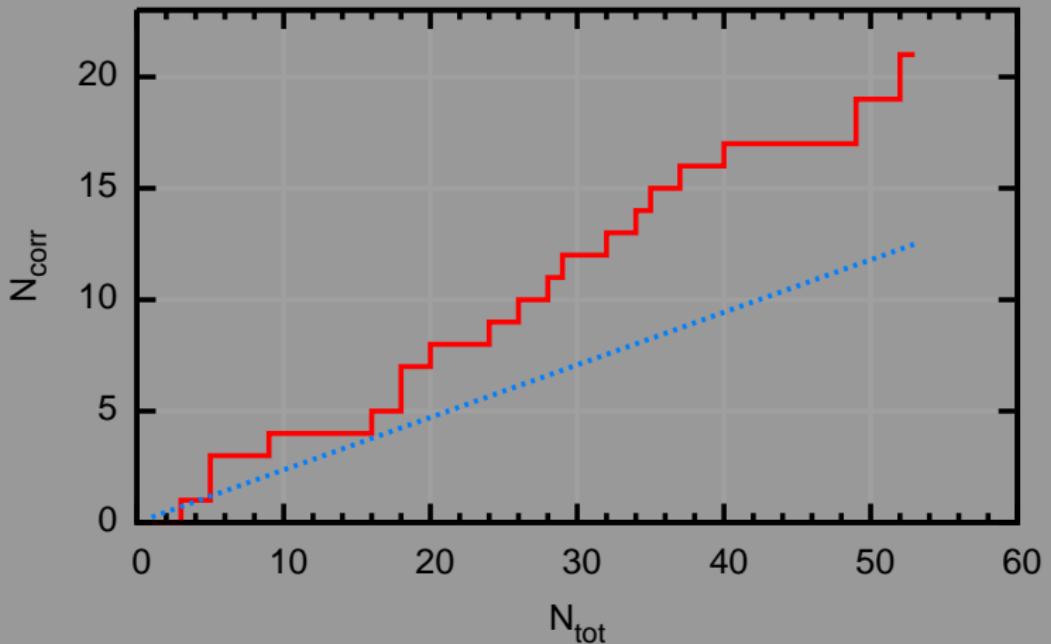
Conclusions

# Correlation with AGN from VCV catalog



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.



Correlating events (red solid) vs. background (blue dotted). Today's values:

$$N_{tot} = 53, \quad N_{cor} = 21, \quad N_{bg} = 12.5$$

$$\text{P-value} = 0.007 (2.7\sigma)$$

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# Correlation with AGN from VCV catalog



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and  
data

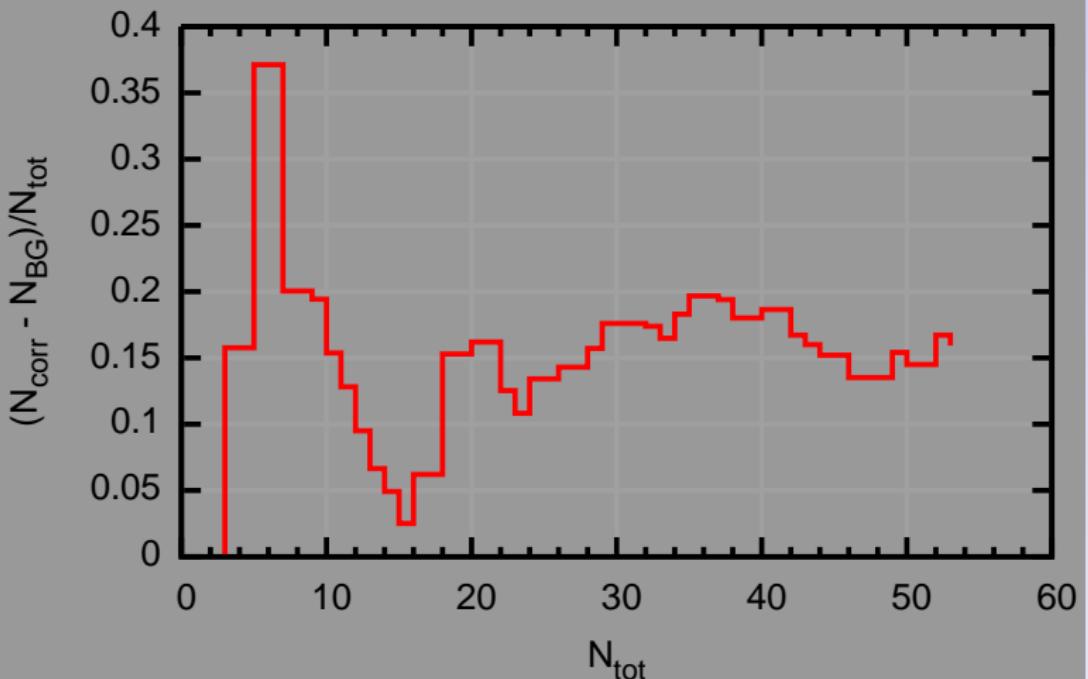
Global  
distributions

Point sources

Hot spot

Other searches

Conclusions



The fractional excess of correlating events  
 $(N_{\text{corr}} - N_{\text{bg}}) / N_{\text{tot}}$ .



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

# HOT SPOT UPDATE

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# HOT SPOT: 6 yr update



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

Global distributions

Point sources

Hot spot

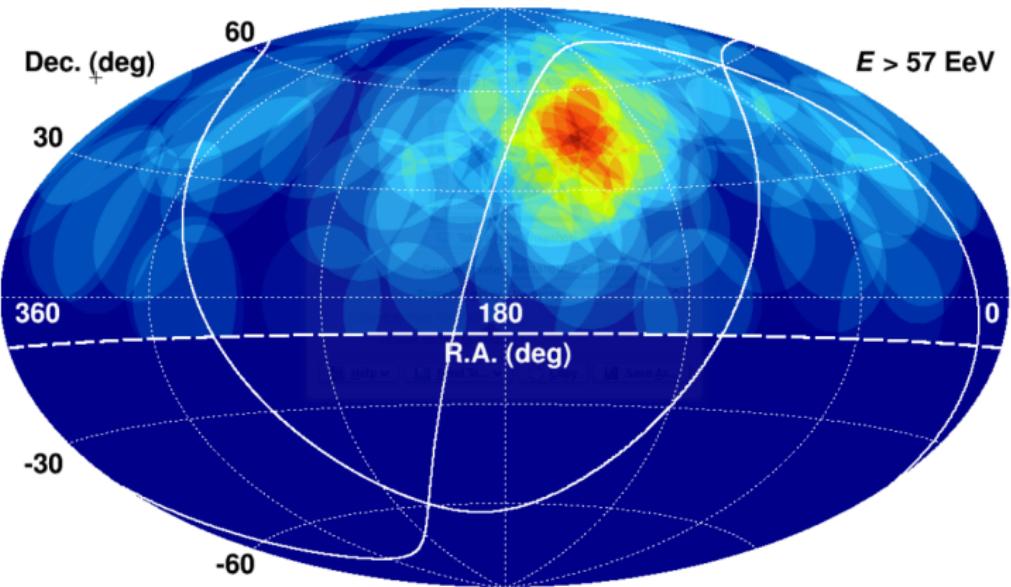
Other searches

Conclusions

- ▶ Reconstruction with loose cuts optimized for statistics (**72** events above 57 EeV in 5 yr).
- ▶ “Hot spot” within the circle of radius  $20^\circ$  centered at RA =  $146^\circ$ , Dec. =  $43^\circ$  [ApJ 790:L21 (2014)].  
Significance of excess  $5.1\sigma$  (pre-trial).
- ▶ After accounting for arbitrary position and opening angles  $15^\circ$ ,  $20^\circ$ ,  $25^\circ$ ,  $30^\circ$ ,  $35^\circ$  the significance  **$3.4\sigma$**  (post-trial).

# HOT SPOT: 6 yr update

## Oversampling with 20°-radius circle



### TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and  
data

Global  
distributions

Point sources

Hot spot

Other searches

Conclusions

# HOT SPOT: 6 yr update



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and  
data

Global  
distributions

Point sources

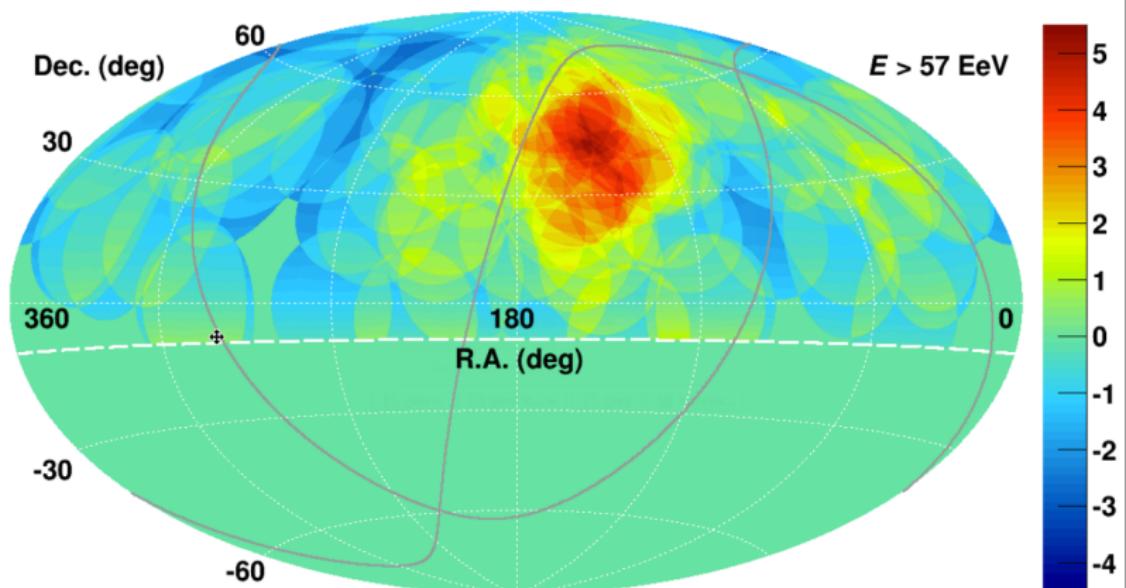
Hot spot

Other searches

Conclusions

## Significance Map (Li-Ma) 6 years

Oversampling with 20°-radius circle



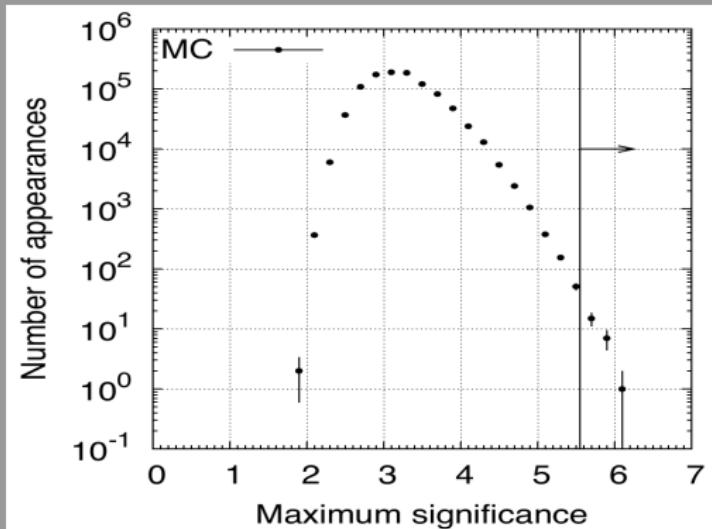
Max significance  $5.55\sigma$  ( $N_{\text{on}} = 23$ ,  $N_{\text{bg}} = 5.49$ ) for 6 years (5.07 $\sigma$  for 5 years)  
Centered at R.A=148.4°, Dec.=44.5° (shifted from SGP by 17°)



# HOT SPOT: 6 yr update

Significance (same procedure as ApJL 790 (2014) L21):

- oversampling at  $15^\circ$ ,  $20^\circ$ ,  $25^\circ$ ,  $30^\circ$ ,  $35^\circ$ , moving center



- Pre-trial:  $P = 5.55\sigma$ ; Post-trial  $P = 3.1 \times 10^{-5}$  ( $4\sigma$ )
- Blind search with 1yr data:  
expected 0.94, observed 3,  $P = 0.07$

## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# HOT SPOT: 6 yr update



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

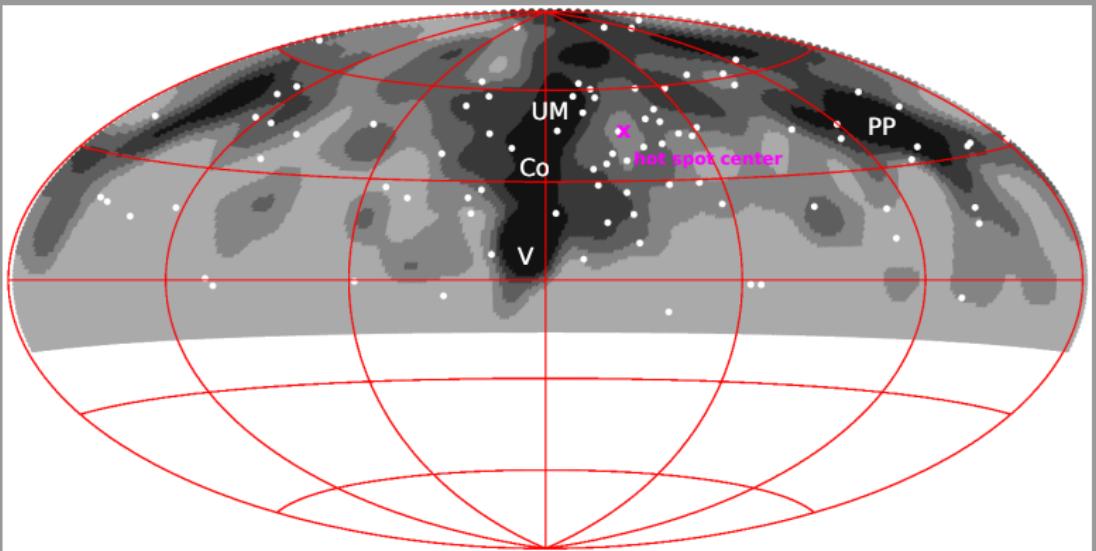
Global distributions

Point sources

Hot spot

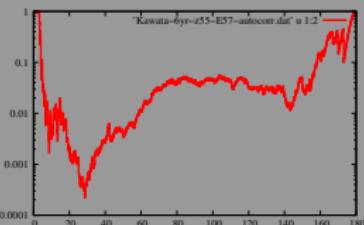
Other searches

Conclusions



Events vs. flux expectation from the LSS (equatorial coordinates). Darker color represents larger flux.

# HOT SPOT: 6 yr update



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

TA detector and data

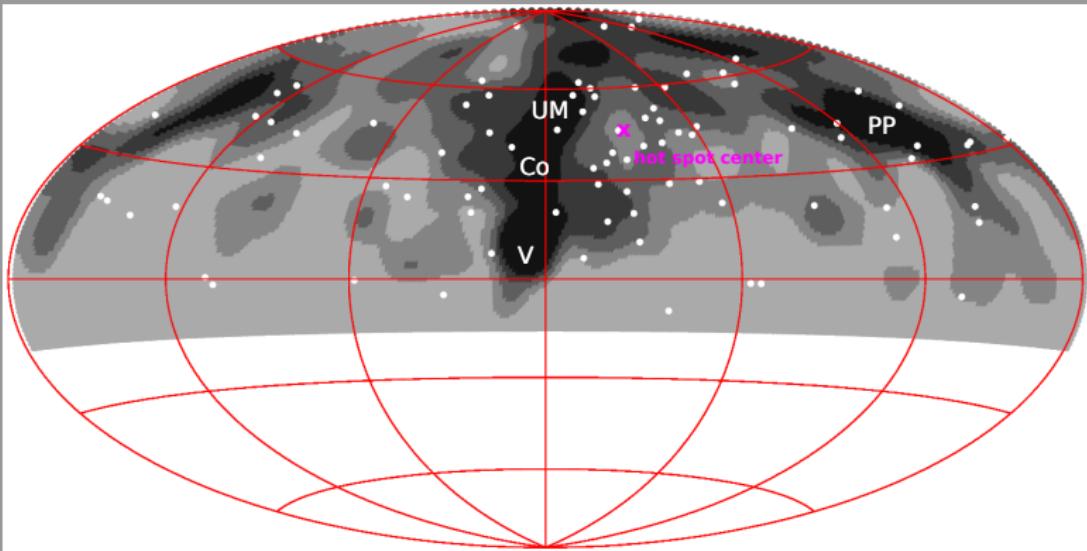
Global distributions

Point sources

Hot spot

Other searches

Conclusions



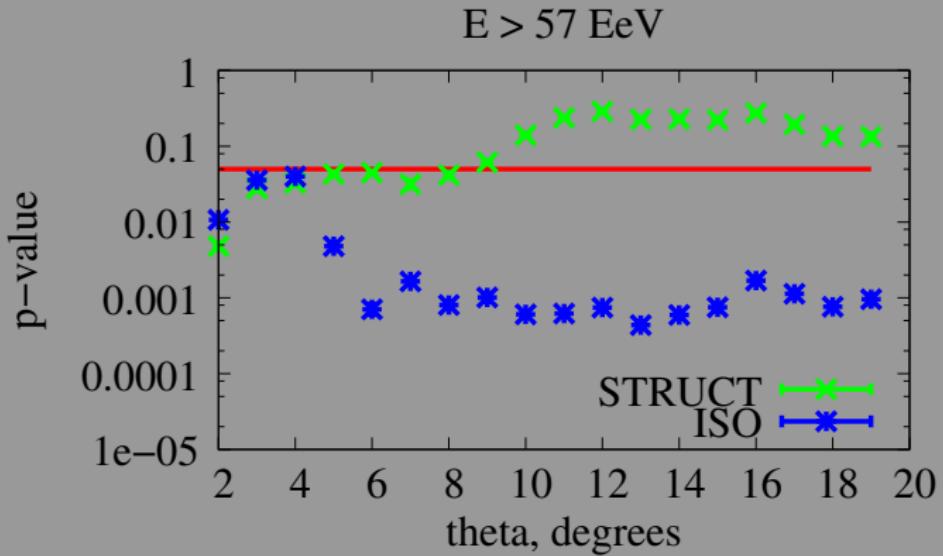
Events vs. flux expectation from the LSS (equatorial coordinates). Darker color represents larger flux.

# Statistical test for compatibility with LSS & isotropy



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.



Compatibility as a function of smearing angle theta (low p-values = incompatibility).



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

# OTHER SEARCHES

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# SEARCHES COMPATIBLE WITH ISOTROPY



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

- ▶ Search for correlations with catalogs of sources detected in different bands [ApJ, 777, 2013, 88]:
  - ▶ 13th VCV AGN catalog
  - ▶ the third Cambridge catalog of radio sources catalog (3CRR)
  - ▶ the 2MASS (the Two Micron All-Sky Survey) redshift survey catalog (2MRS)
  - ▶ Swift BAT 58-Month hard X-ray survey catalog
  - ▶ 2nd Fermi AGN catalog (2LAC)
- ▶ Low energy  $\sim 10^{18}$  eV point-like source search [arXiv:1407.6145]
- ▶ Medium-scale anisotropy at low energy  $\sim 10^{18}$  eV
- ▶ Correlations with LSS at 10 EeV and 40 EeV
- ▶ Harmonic analysis [arXiv:1409.3128]

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions

# CONCLUSIONS



## TA ANISOTROPY SUMMARY

P. Tinyakov  
for the Telescope  
Array  
Collaboration.

- ▶ Isotropy at low energies
- ▶ “Hot spot” of a medium angular scale at highest energies  $E > 57$  EeV; current significance  $\sim 4\sigma$  (post-trial). Shows up in various other tests as incompatibility with isotropy:
  - ▶ distribution in RA:  $p \sim 0.02$
  - ▶ autocorrelation function:  $\sim 0.001$  at  $\delta \sim 20^\circ - 25^\circ$
  - ▶ correlation with AGN:  $p \sim 0.01$
  - ▶ correlation with LSS:  $p \sim 0.001$
- ▶  $\Rightarrow$  Should be tested with a few times larger statistics

TA detector and data

Global distributions

Point sources

Hot spot

Other searches

Conclusions