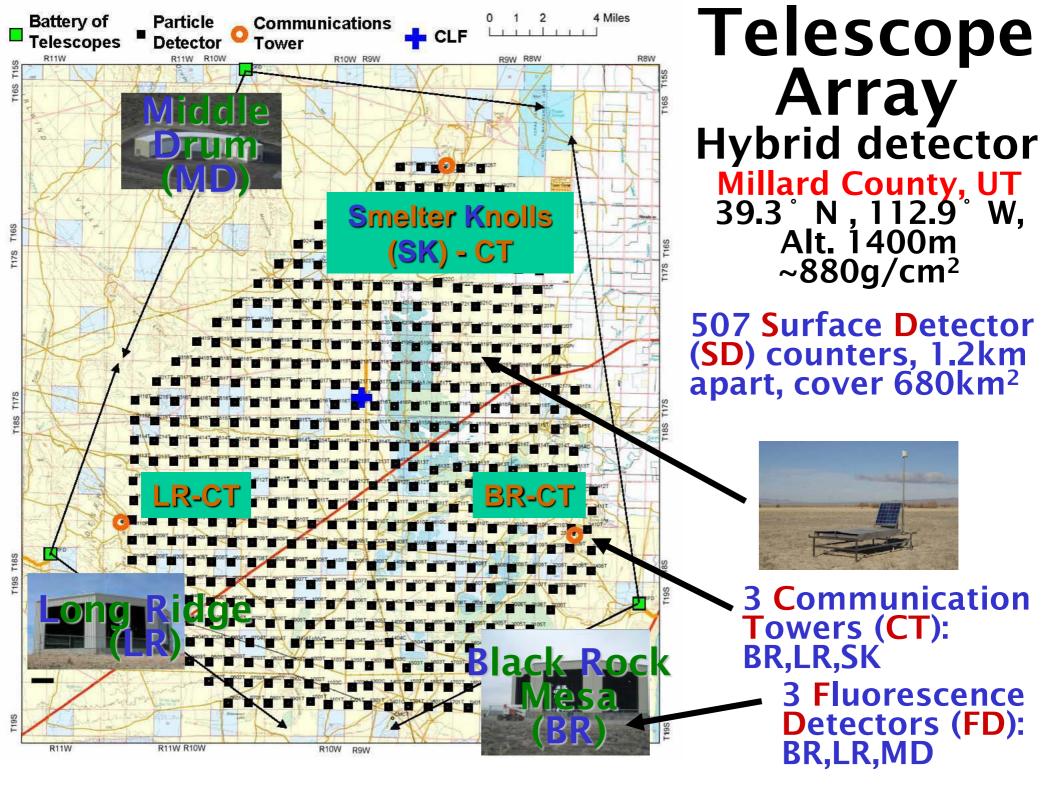
High Energy Spectrum Measured by the Telescope Array Experiment



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Outline

- Telescope Array Experiment
- Surface Detector Spectrum
- Fluorescence Mono and Hybrid Spectra
- Comparison with HiRes and Auger
- Conclusions



TA FD

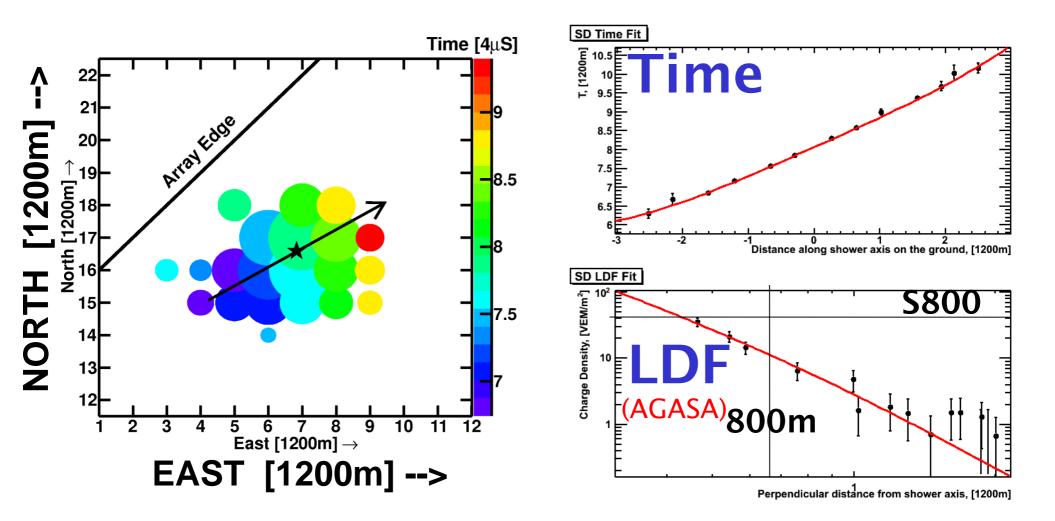


TA Surface Detector

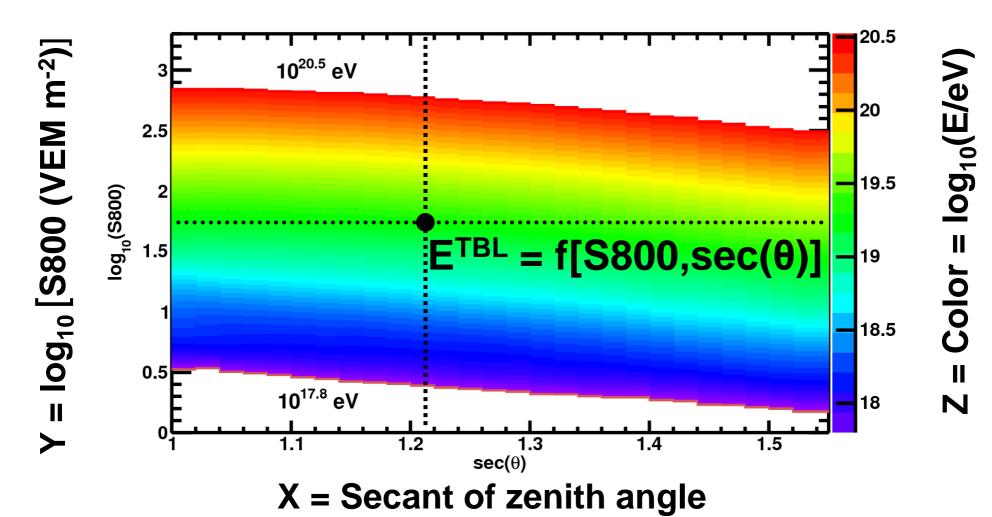
- Powered by solar cells; radio readout.
- In operation since March, 2008.
- Self-calibration using single muons.
- Energy deposited by cosmic ray shower particles is measured in VEM units (Vertical Equivalent Muon = energy deposited by a vertical minimum ionizing muon)



Surface Detector Event

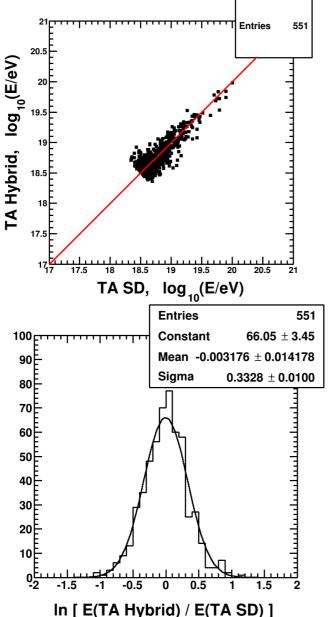


SD Energy 1/2



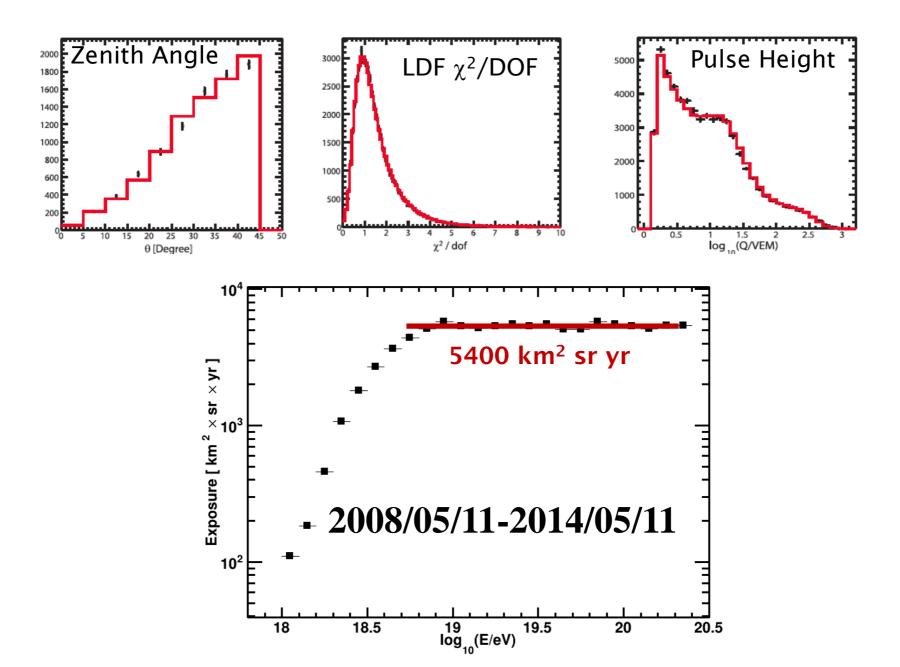
- A look-up table made from the Monte-Carlo
- Event energy (E^{TBL}) = function of *reconstructed* S800 and sec(θ)
- Energy reconstruction $\leftarrow \rightarrow$ interpolation between S800 vs sec(θ) contours of constant values of E^{TBL}
- The overall energy scale locked to the fluorescence detector

SD Energy 2/2: Energy Scale Set

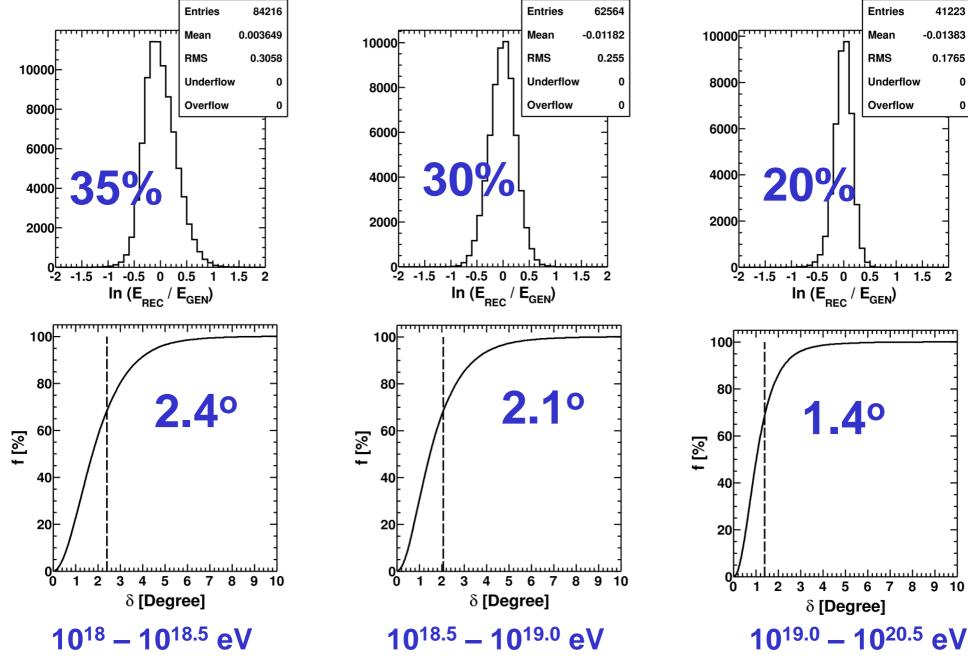


- Energy scale locked to the FD to reduce the systematic due to the model
- Use events well reconstructed separately by SD and FD in hybrid mode:
- $SD \cap [BR U LR U MD Hybrid]$ $E^{FINAL} = E^{TBL} / 1.27$
- TOP figure: E^{FINAL} vs E^{FD} scatter plot
- BOTTOM figure: histogram of E^{FINAL} / E^{FD} ratio
- 2008/05/11-2013/05/04

Exposure from Monte Carlo

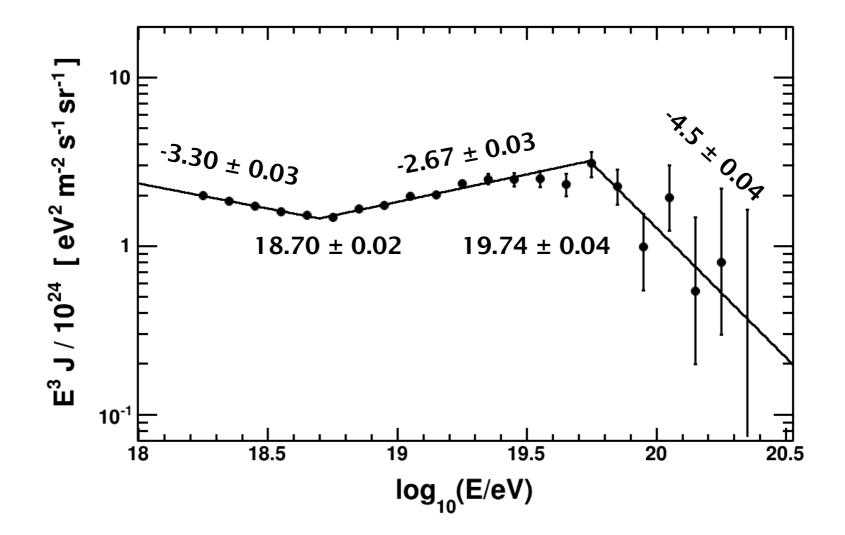


SD Resolution from Monte Carlo

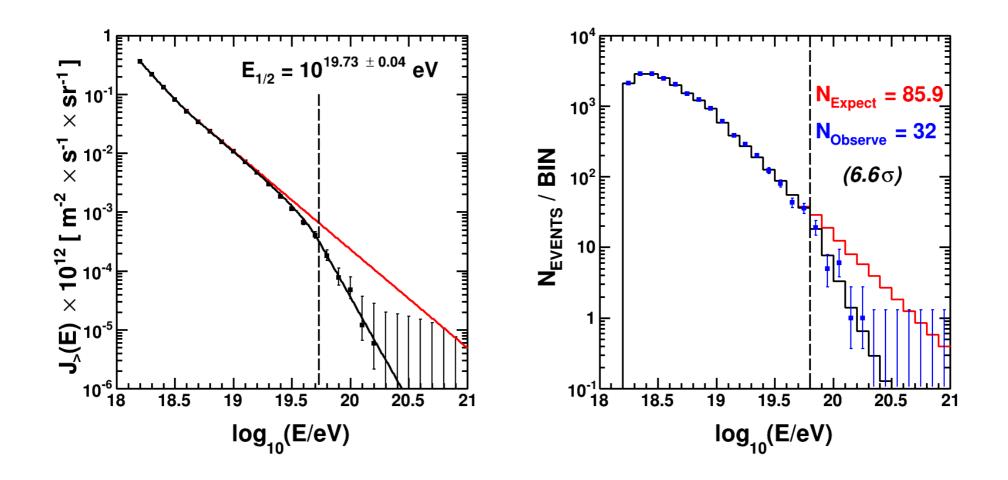


SD Spectrum

2008/05/11-2014/05/11



GZK Cutoff



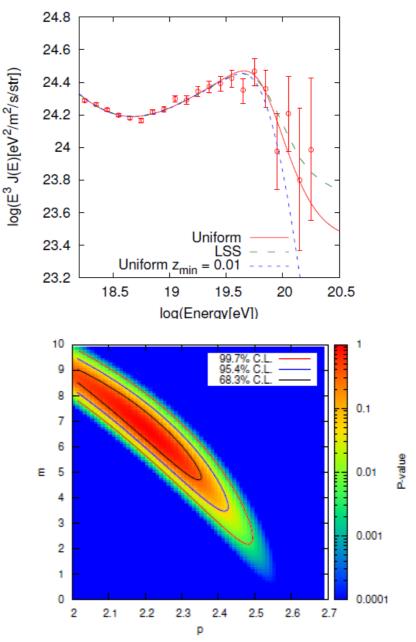
Fit spectrum to energy-loss model

Inputs:

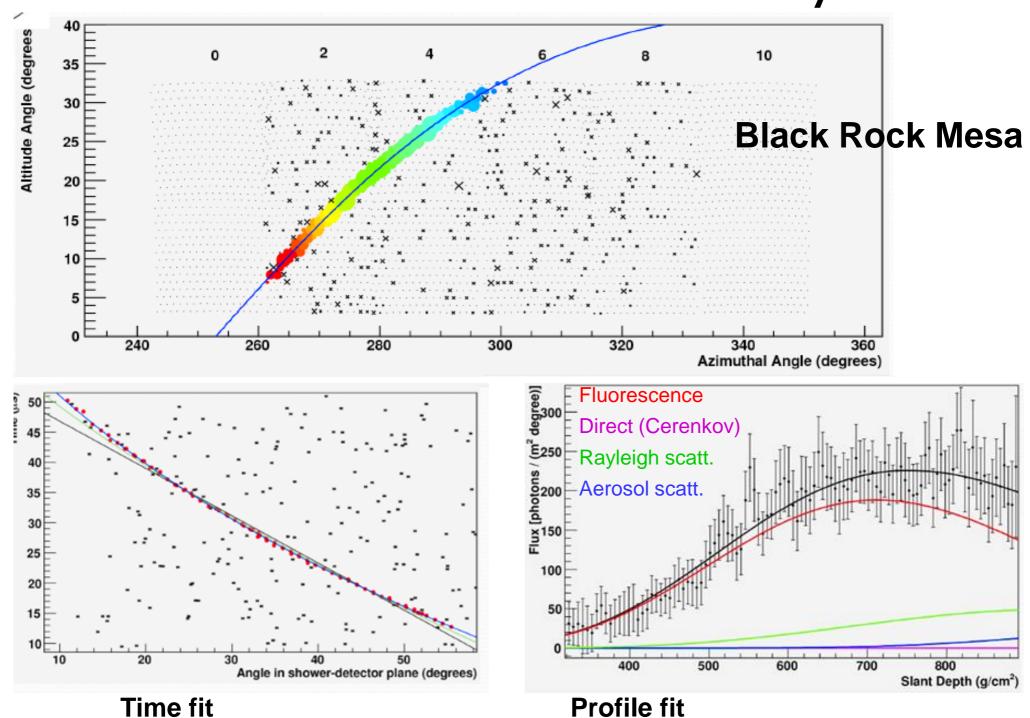
- Pion photoproduction and e+e- pair production;
- 2. Hubble expansion.

Fitting parameters:

- 1. Power law at the source, E^{-p}
- 2. Evolution of the sources, (1+z)^m

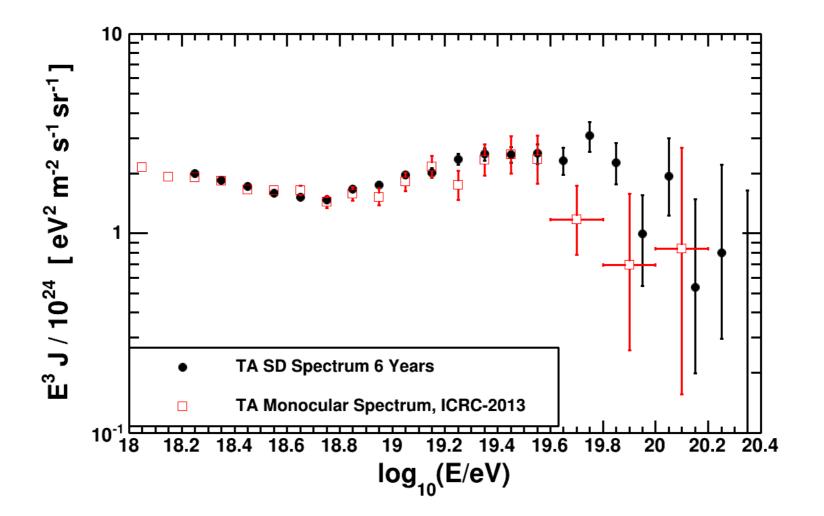


Fluorescence Mono Analysis



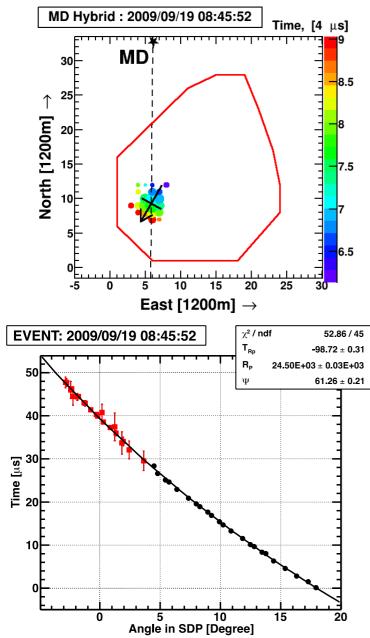
Time fit

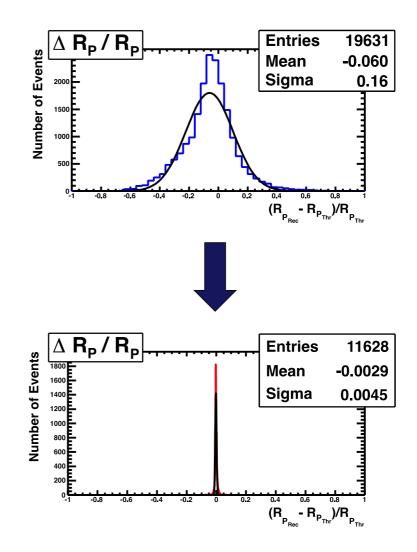
Combined Fluorescence Mono



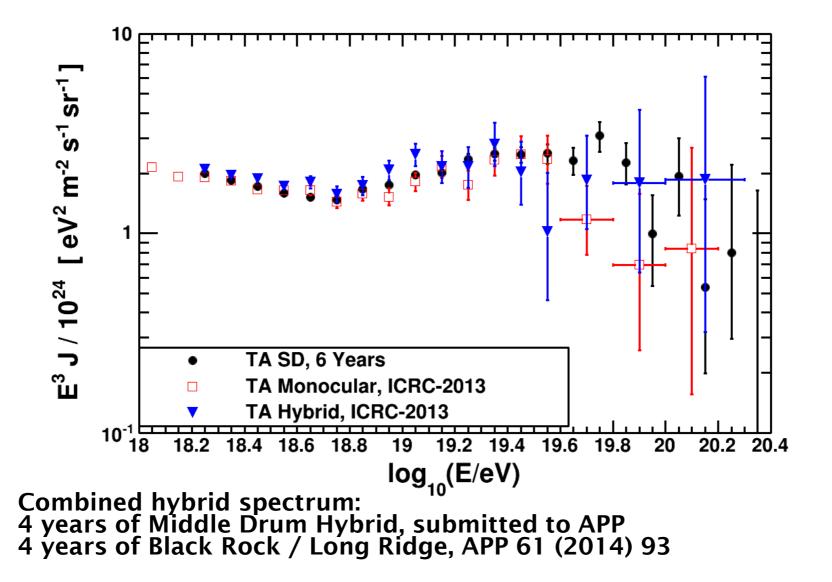
Combined mono spectrum: 3 years of Middle Drum, APP 39 (2012) 109 3.5 years of Black Rock / Long Ridge, APP 48 (2013) 16 Use geometric mean of energies, account for the overlaps in the exposure

Hybrid analysis





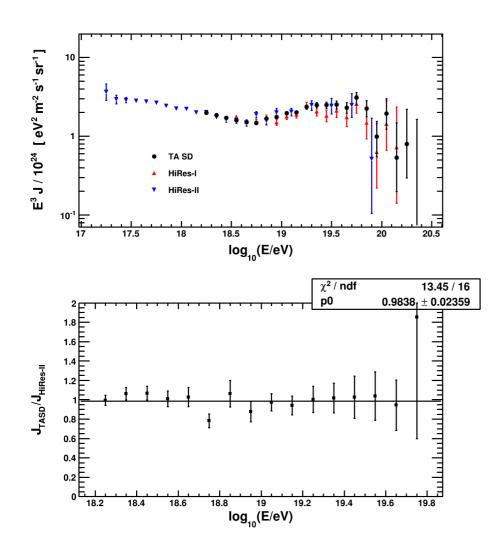
Combined Hybrid Spectrum



Next

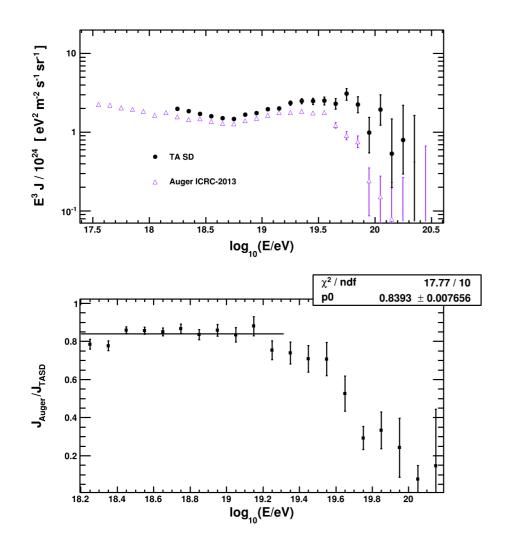
Compare with other experiments: HiRes-II, HiRes-II (PRL-2008) Pierre Auger combined (ICRC-2013)

TA SD and HiRes



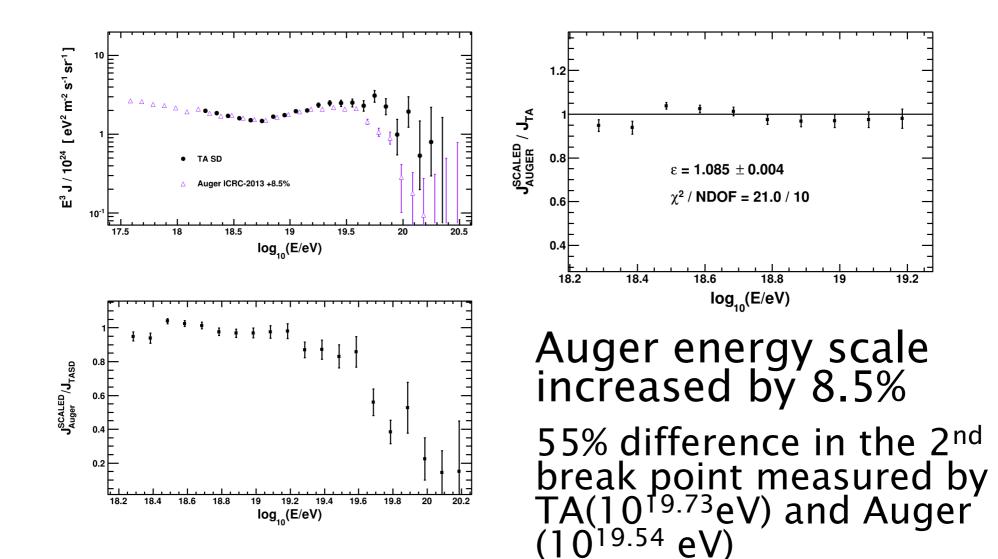
- Different experiments using different techniques
- Looking at the same sky
- Agreement in overall shape
- Same energy scale

TA SD and Pierre Auger

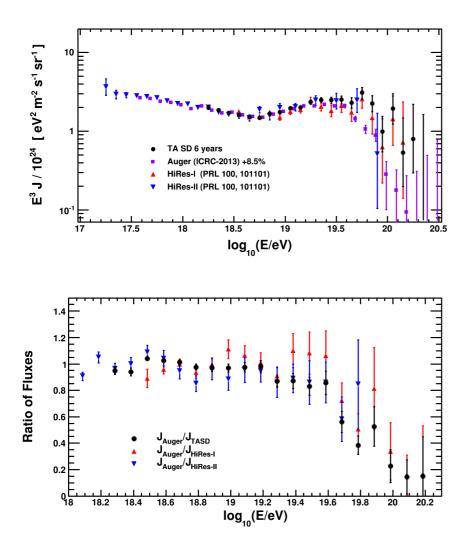


- Similar experiments but looking at different parts of sky
- Auger flux 16% lower than TA in the flat ratio region
- Large differences at the highest energies

TA SD and Pierre Auger

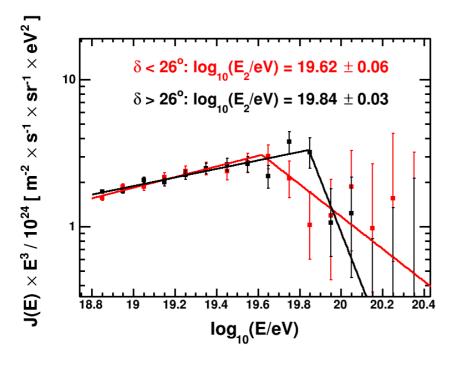


TA SD, Pierre Auger, and HiRes



 Similar feature seen at the highest energies when Auger spectrum is compared to HiRes

Declination dependence of TA spectrum



- Auger and TA sensitive to different parts of the sky
 - TA: delta > -16° (zenith angle < 55)
 - Auger: delta < 26°
- Hint for the declination dependence in TA data (~3σ)

Conclusion

- SD Spectrum updated to 6 years of data, 2008/05/11-2014/05/11
- Energy scale result is unchanged after including more hybrid events into comparison
- TA Mono, Hybrid, and SD spectra all agree
- Fits the simple proton propagation model
- Agree with HiRes, discrepancy with Auger at the highest energies
- Hint of the declination dependence of the second break point

Thank You





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